



## SToRM SAR

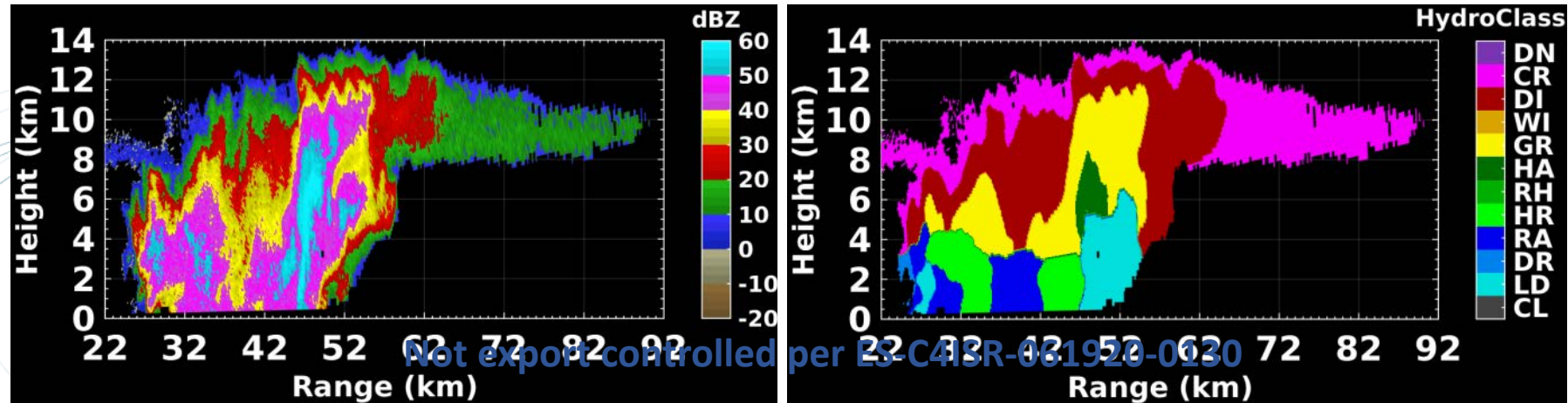
Presenter: Kevin Maschhoff

PI: Kevin Maschhoff, BAE Systems

Team Members: M. Ryba, V. Chandrasakar (CSU)

Program: IIP 19-0012

# What Spatial Resolving Power Is Needed to Observe Intense Mid-Latitude Storms From Space?

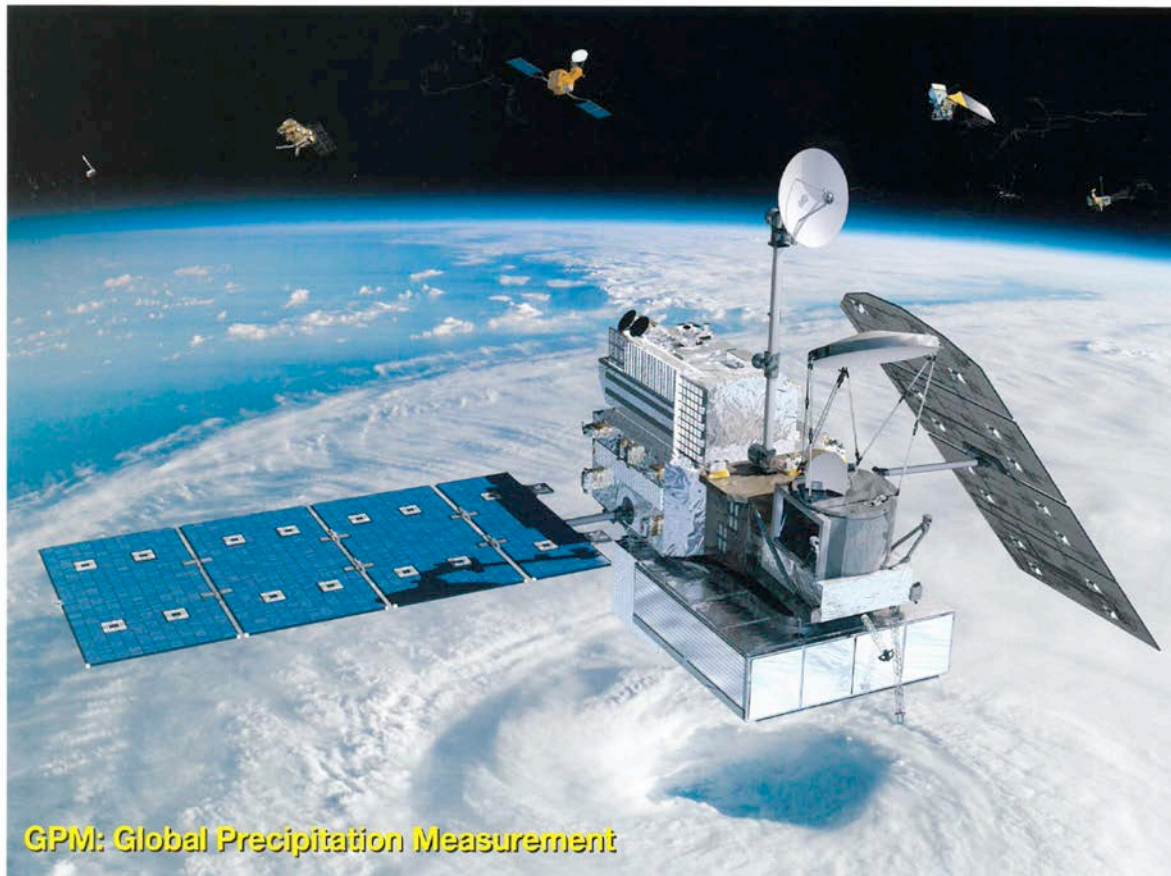


*Radar Reflectivity Cross Section and Hydrometeor (classified) Cross-Section for a Precipitation Event Observed at 23:43UTC, May 29, 2013 by NPOL during the IFLOOD Field Campaign (HydroClass: CL: Clear Air, LD: large drops, DR: drizzle, RA: rain, HR: heavy rain, RH: rain plus hail, HA: hail, GR: graupel, WI: wet ice, DI: dry ice, CR: crystals)*

Ground-Based Radar Observations Illustrate that ~ 1 km Horizontal Resolution Needed to Characterize Intense Storms

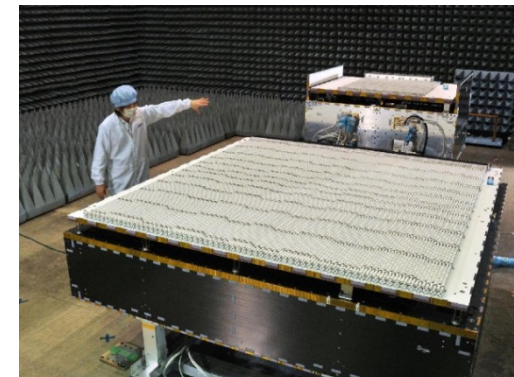
# Problem to Solve-Providing 1 km Horizontal Resolution Precipitation Observations from Space

National Aeronautics and Space Administration

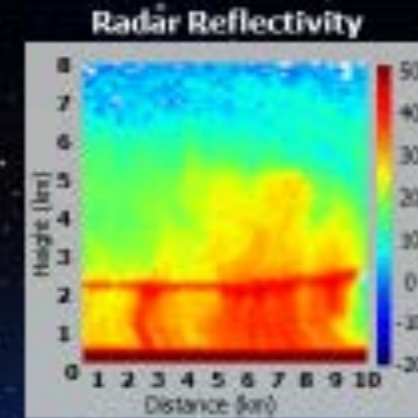


NASA/JAXA's Premier Space-Based Precipitation Radar:

- ~\$1B Mission, a Decade in Development
- Its 5 km Spatial Resolution at Ku-Band is 5x Coarser than Needed For Many Intense Storms



# SToRM - Remote sensing of precipitation using multi-static radar interferometry & tomography



Receiver satellite

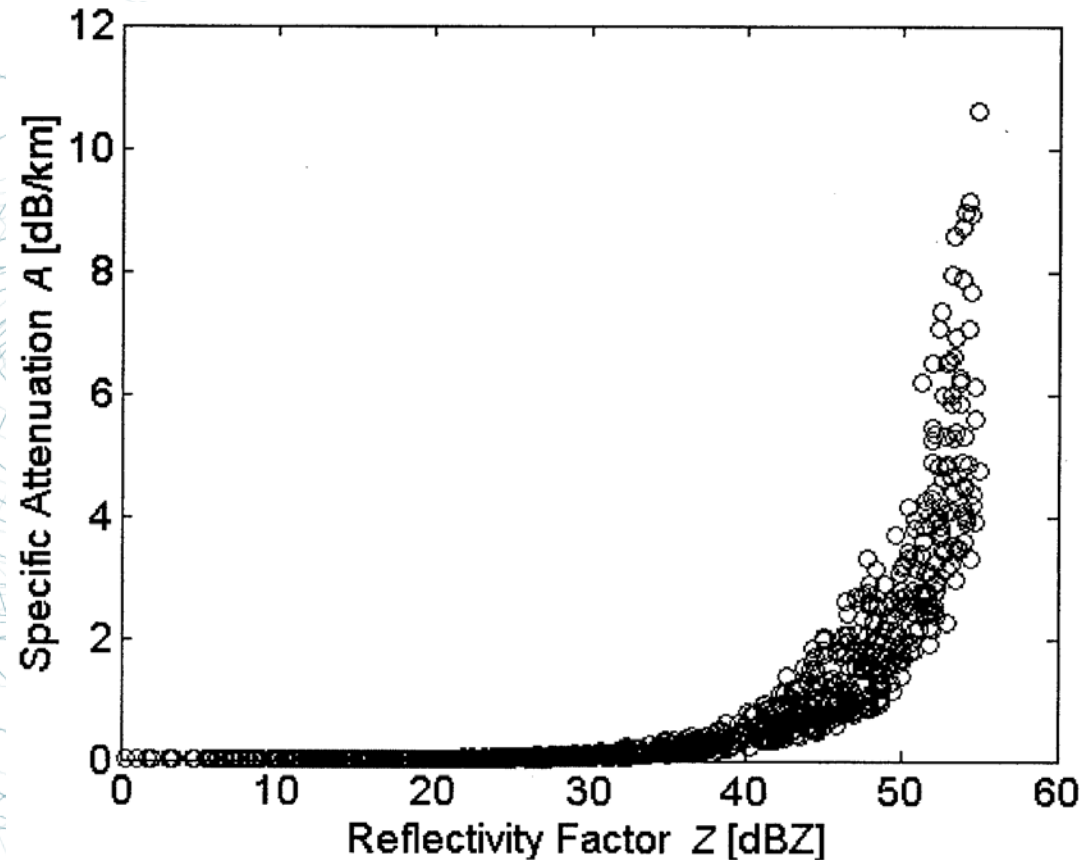
Receiver satellite

Transmitter satellite

## Purpose and Concept Summary

- Precipitation Observations from Space at  $\sim 1$  km horizontal resolution are needed to characterize severe storm processes for weather research
- Providing these observations via up-scaled, larger aperture implementation using current methods would be very costly
- SToRM SAR employs a distributed architecture of micro-satellites, closely synchronized, to implement a networked-radar approach this observation

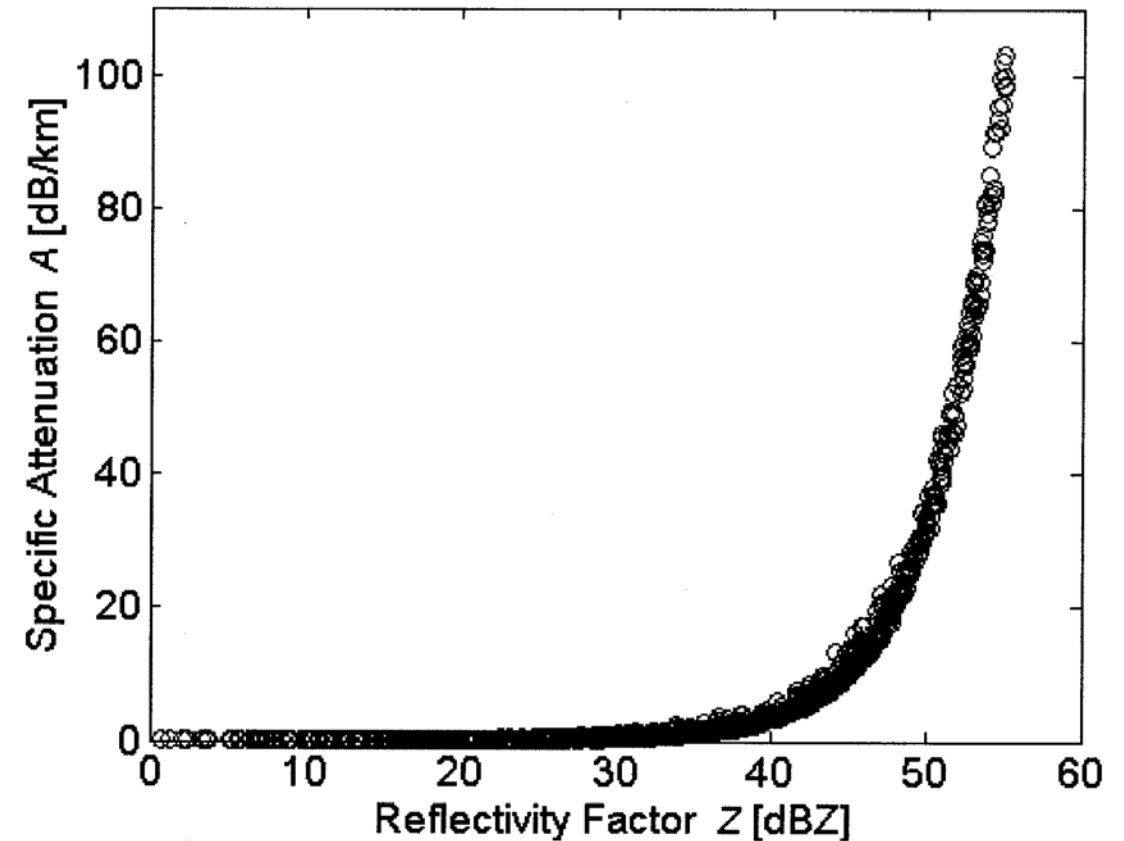
# Specific Attenuation at Ka and Higher Frequencies Limits Vertical Profiling of Intense Convective Storms



Global Mapping of Attenuation at Ku- and Ka-Band  
V. Chandrasekar, Hiroki Fukatsu, and K. Mubarak

IEEE TRANSACTIONS ON GEOSCIENCE AND REMOTE SENSING, VOL. 41, NO. 10, OCTOBER 2003

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While Real-Aperture Diffraction-limited Footprints (and Antennas) are Smaller at Ka, Path Attenuation is Much Larger

July 9, 2020

# SToRM Multi-Static/SpotLight Observation Geometry

## Key Multi-Static Observation Variables

Observation Parameter	Symbol
Transmitter → Reference Receiver Separation	B
Receiver → Receiver Separation	b
Orbital Height	H
Altitude of Observed Field	A
Cross-Track Position (relative to Transmitter)	$\zeta$
Cross-Track Position-Relative to Tx-Rcv <sub>1</sub> Specular Point	$\Delta\zeta$
Range Difference—The Difference in Distance between Observation Field Element and the Receivers	$\Delta R$

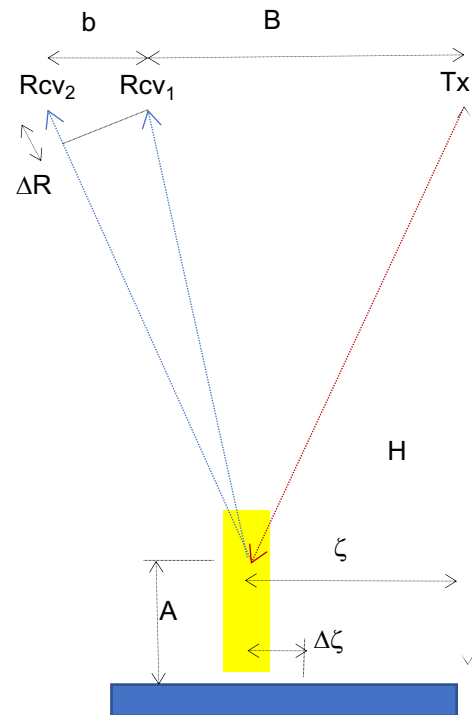
$$R_1^2 = (H - A)^2 + (B/2 - \Delta\zeta)^2$$

$$R_2^2 = (H - A)^2 + (\frac{B}{2} + b - \Delta\zeta)^2$$

$$\Delta R = R_2 - R_1$$

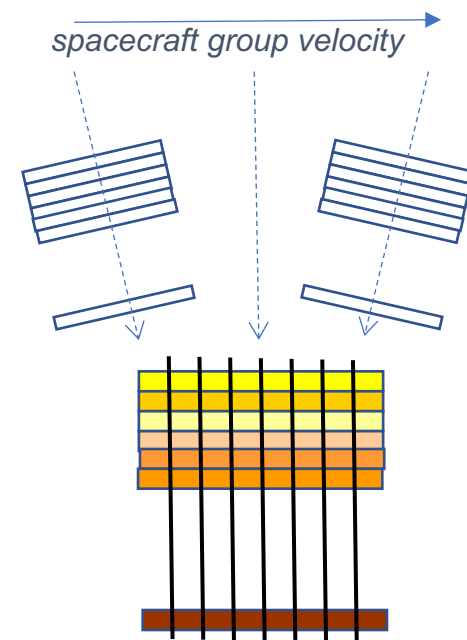
For Large R,  
 $\Delta R \approx \text{const.} \times \Delta\zeta$

## Cross-Track Observation



*For Typical LEO Orbital Heights,  $\Delta R$  Varies Linearly with Cross-range Position, and is Largely Independent of Altitude, A*

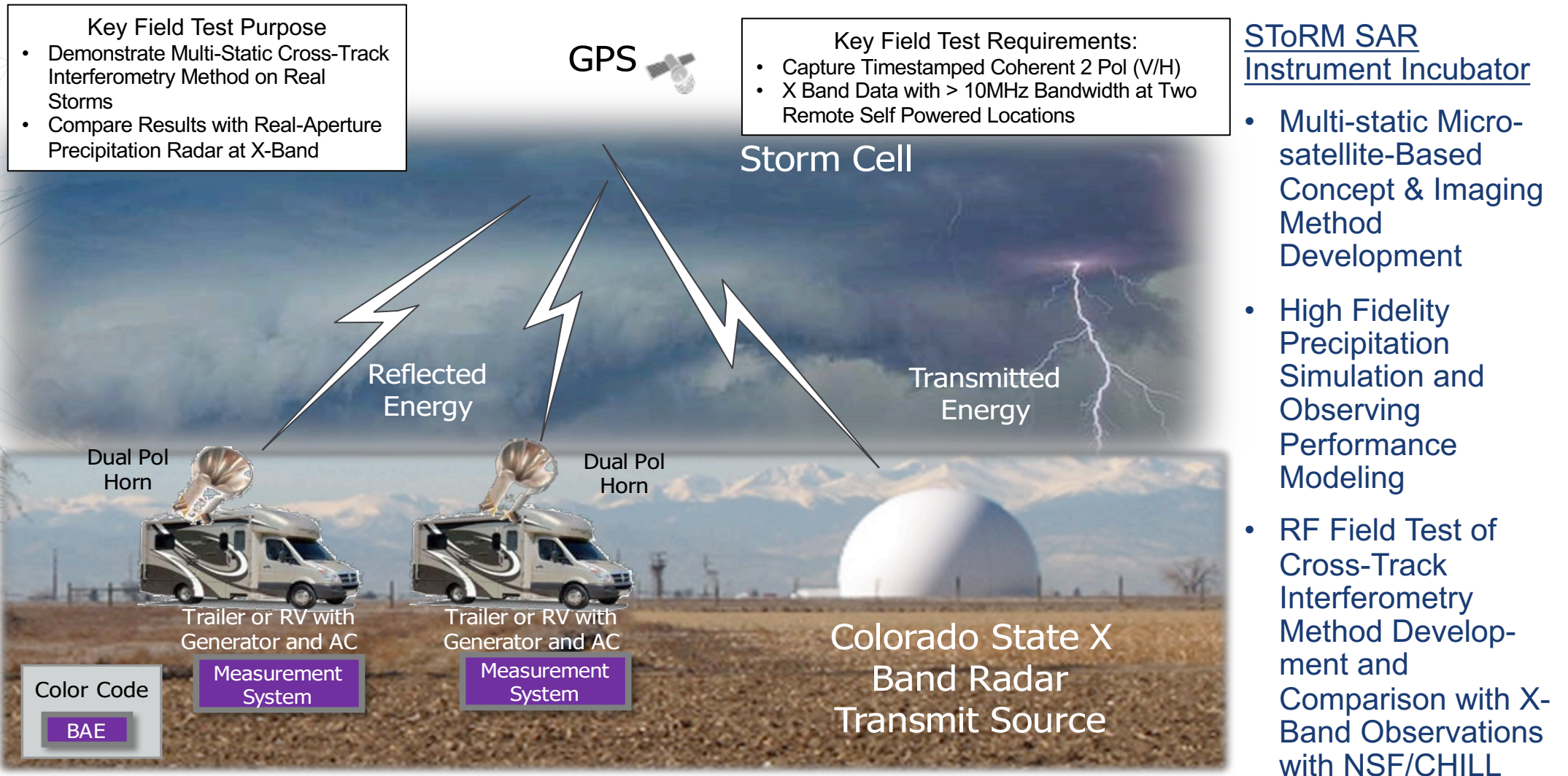
## Along-Track Observation



*Employing SAR Spotlight-Mode Methods, Layered Structure is Seen over a Range of Angles-Enabling a Form of Tomography*

# Next Steps

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